



DA SUNANKA
RABBI NAKA
FARA KOMAI



Numerical Moduli Stabilisation towards Calabi-Yau Data Exploration

Shehu AbdusSalam

Department of Physics, Shahid Beheshti University, Tehran

in collaboration with:

S.Abel, M.Cicoli, F.Quevedo and P.Shukla

SUSY-2019, Texas A&M University – Corpus Christi, Texas

May 22, 2019



- ▶ **Calabi-Yau** data (CYD) \rightarrow 1411.1418
Aim: phenomenology-based classification
- ▶ **How?** Bayes factors & Jeffreys' scale w.r.t. **LVS AdS**
Via: Multi modular fields stabilisation & nested sampling
- ▶ **Outlook:** Machine-learn the Calabi-Yau data,
Guide: phenomenology, stringy remnants \rightarrow 1801.03503
- ▶ **This talk:** P_{11169} AdS vs dS scenarios comparison
Based on: explorations towards minimal V_0

4D Compactification

+ fluxes + non-perturbative effects + α' -corrections

generates super, Kähler, and $N = 1$, $D = 4$ scalar potentials

$$W = W_0 + A_i e^{-a_i T_i}, \quad i = 1, 2$$

$$K = -2 \log \left(\mathcal{V} - \frac{\xi}{2} \right) + 2 \log g_s + K_{CS},$$

$$V = e^K \left[K^{i\bar{j}} D_i W D_{\bar{j}} \bar{W} - 3 |W|^2 \right]$$

where $D_i W = \partial_i W + (\partial_i K) W$, $K^{i\bar{j}} = (\partial_i \partial_{\bar{j}} K)^{-1}$

LVS: two moduli stabilised & break SUSY at large volume

Parameters, θ : $\{g_s, W_0, a_i, A_i\}$

Possible Scenarios: $\mathcal{H}_0 \equiv$ **Anti-de Sitter** or

$\mathcal{H}_1 \equiv$ **de Sitter (without upliftment)**

Numerical Moduli Stabilisation



Given θ , Find T_i at (deepest) minimum of $V(\theta, T_i)$
Genetic or Dlib C++ Library minimisation algorithm

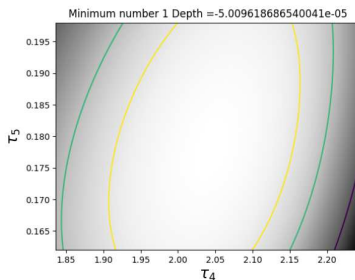


Figure: Deepest of 6 minima found using Genetic algorithm. A check for Dlib minimiser used.

Apply towards Calabi-Yau data pheno-based **classification**

► **Bayes: evidence**, $Z \times \text{posterior} = \text{prior} \times \text{likelihood}$

$$p(d|\mathcal{H}) \times p(\theta|d, \mathcal{H}) = p(\theta|\mathcal{H}) \times p(d|\theta, \mathcal{H})$$

$$Z = p(d|\mathcal{H}) = \int p(d|\theta, \mathcal{H})p(\theta|\mathcal{H}) d\theta \quad [\text{MultiNest}]$$

Consider: $d \equiv \text{cosmological constant}(\Lambda)$ is very small

Likelihood function, $\mathcal{L}(\theta) = p(d|\theta, \mathcal{H})$

With V_0 from numerical moduli stabilisations, get Z using

$$\log \mathcal{L} \sim -|V_0| \quad \text{or, alternatively,} \quad \mathcal{L} \sim \exp^{-(V_0-\Lambda)^2}$$

- ▶ **Comparison:** Bayes factor, $K = \frac{z_0}{z_1}$
- ▶ **Jeffreys' scale:**

Bayes factors, K	Comparison Remarks	Classification
1 to 3.2	Inconclusive	1
3.2 to 10	Weak	2
10 to 100	Moderate	3
> 100	Decisive	4

Table: Jeffreys' scale for the interpretation of Bayes factors.

Classification, say w.r.t. LVS AdS. **0**, comparison not possible. **5**, **6** or **7**, compared geometry wins weakly, moderately or decisively respectively.

LVS de Sitter vs Anti-de Sitter [preliminary]



P_{11169} parameters, θ	Range
g_s	$10^{-3} - 0.3$
W_0	$10^{-11} - 10^2$
$a_{1,2}$	$2\pi/10 - 2\pi$
$A_{1,2}$	$10^{-1} - 10^1$

Scenario	$\log_e Z$	K	Class
AdS	-0.366	-	reference hypothesis
dS	-3×10^{-14}	3×10^3	7

de Sitter vacua are obtainable without need for uplift terms

LVS de Sitter vs Anti-de Sitter [preliminary]

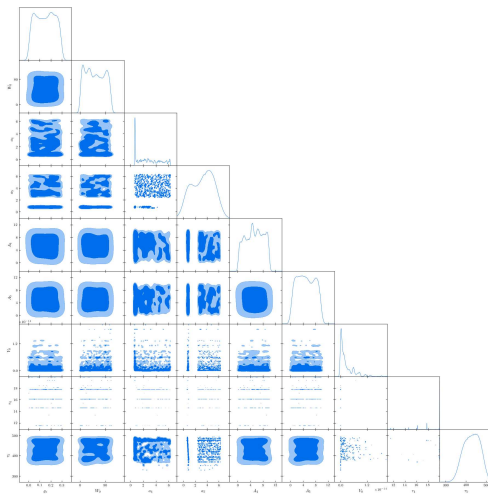


Figure: de Sitter scenario.



- ▶ **Demonstrated:**
Multi moduli stabilisation is within reach [dedicated to Joe Polchinski]
- ▶ **Proposed:** reference point model/CY. Only Λ used.
Extend to include inflaton or particle physics constraints
- ▶ **Classification:** Bayes factors-based. Something else ?
- ▶ **de Sitter Vacua:** found many, without uplifting terms.



Thanks for Listening!